Radiative corrections to anti-neutrino proton scattering at low energies

F. Myhrer

University of South Carolina

Collaborators:

K. Kubodera and U. Raha

MOTIVATION Neutrino oscillations:

- Atmospheric neutrinos high energies
- Solar neutrinos —— low energies ($E_{\nu}^{max} < 20 \text{ MeV}$)
- Reactor neutrinos very low energies: KAMLAND, Chooz, Daya Bay, RENO

Anti-neutrinos from nuclear reactors are ideal in order to determine the neutrino mixing parameter $\sin^2(2\Theta_{13})$.

The mixing angle Θ_{13} is the last one to be determined. It is non-zero. Experimental values published in Phys. Rev. Letters 108, 2012.

The reactor anti-neutrinos hit a H_2 target. Measure the positron yield Y_{e^+} for a given reactor $\bar{\nu}_e$ flux $n_{\nu}(E_{\nu})$: $Y_{e^+} = n_{\nu}(E_{\nu}) \ \sigma(\bar{\nu}_e + p \to e^+ + n)$

NEUTRINO OSCILLATIONS

Neutrino mixing proposed by
Pontecorvo, Sov. Phys. JEPT bf 6,429 (1957) and 26, 984 (1968)
and
Maki, Nakagawa and Sakata, Prog. Theor. Phys. 28, 870 (1962)

Three angles and one phase determine the mixing of the three neutrino species.

The weak interactions act on the three neutrino flavors ν_e, ν_μ and ν_τ .

When Θ_{13} is "large" one could expect to be able to measure the leptonic CP-violating phase

$$V_{13} \propto \sin\Theta_{13} e^{-i\delta_{CP}}$$

An accurate evaluation of $\sigma(\bar{\nu}_e + p \rightarrow e^+ + n)$, requires knowledge of the RADIATIVE and nucleon RECOIL corrections.

Effective Field Theory

allows a <u>systematic</u> evaluation of these corrections within a <u>unified approach</u> to <u>electro-weak</u> and <u>strong</u> processes at <u>low energies</u>. in a gauge-invariant and model-independent way.

- Long wavelength probes are <u>insensitive</u> to short distance details.
- Low-energy-constants (LECs) parametrize short-distance physics.
- The effective degrees of freedom are hadrons.
- The interactions among hadrons reflect the <u>symmetries</u> and <u>symmetry-breakings</u> of QCD.

We use Heavy Baryon Chiral Perturbation Theory (HBChPT)

Baryons treated non-relativistically.

Leptons treated relativistically.

The expansion parameters: $\left(\frac{Q}{\Lambda_\chi}\right) \simeq \left(\frac{\alpha}{2\pi}\right) \sim 10^{-3}; \; \Lambda_\chi \sim 1 \; {\rm GeV}.$

"The Lowest Order (LO)" lagrangian:

$$\mathcal{L}_{QED} = -\frac{1}{4} F^{\mu\nu} F_{\mu\nu} + \bar{\psi}_e (i\gamma \cdot D) \psi_e - m_e \bar{\psi}_e \psi_e + \bar{\psi}_\nu (i\gamma \cdot \partial) \psi_\nu,$$

$$\mathcal{L}_{NN\gamma} = \bar{N} (iv \cdot D) N,$$

$$\mathcal{L}_{LO} = -\left(\frac{G_F V_{ud}}{\sqrt{2}}\right) \{\bar{\psi}_e \gamma_\mu (1 - \gamma_5) \psi_\nu\} \times \{\bar{N} \tau^+ [v^\mu - 2 g_A S^\mu] N\}.$$

The lagrangian contains several LECs assumed known from other processes, e.g. the Fermi constant G_F and axial coupling constant g_A .

The "Next to Leading Order (NLO)" interaction lagrangian includes nucleon recoil:

$$\mathcal{L}_{NLO} = \left(\frac{\alpha}{4\pi} e_{1}\right) \bar{\psi}_{e} \left(i\gamma \cdot \partial\right) \psi_{e} + \left(\frac{\alpha}{8\pi} e_{2}\right) \bar{N} (1 + \tau_{3}) \left(iv \cdot \partial\right) N + \\ - \frac{(G_{F} V_{ud})}{\sqrt{2}} \left\{ \bar{\psi}_{e} \gamma_{\mu} (1 - \gamma_{5}) \psi_{\nu} \right\} \times \left\{ \bar{N} \tau^{+} \left[\left(\frac{\alpha}{4\pi} e_{V}\right) v^{\mu} - 2 g_{A} \left(\frac{\alpha}{4\pi} e_{A}\right) S^{\mu} \right] N + \\ + \left(\frac{1}{2m_{N}}\right) \left\{ \bar{N} \tau^{+} \left[i (v^{\mu} v^{\nu} - g^{\mu\nu}) (\overleftarrow{\partial} - \overrightarrow{\partial})_{\nu} - 2i \mu_{V} [S^{\mu}, S \cdot (\overleftarrow{\partial} + \overrightarrow{\partial})] \right. \\ \left. - 2i g_{A} v^{\mu} S \cdot (\overleftarrow{\partial} - \overrightarrow{\partial}) \right] N \right\} \right\}$$

Only one linear combination of the radiative LECs is relevant:

$$e_V^{(R)}(\mu = m_N) = e_V - \frac{1}{2}(e_1 + e_2) + \frac{3}{2}\left[\frac{2}{4-d} - \gamma_E + \ln(4\pi) + 1\right]$$

The LEC μ_V is the nucleon isovector magnetic moment.

The lowest order differential cross section:

$$\frac{\mathrm{d}\sigma_{LO}}{\mathrm{d}(\cos\theta_e)} = \left(\frac{G_F V_{ud}}{\sqrt{2}}\right)^2 f_{LO}(E,\beta) \left[(1+3g_A^2) + (1-g_A^2)\beta\cos\theta_e \right]$$

Positron velocity: $\beta = \sqrt{E^2 - m_e^2}/E$

Phase-space factor: $f_{LO}(E,\beta) = \beta E^2/\pi$

NLO differential cross section:

$$\frac{\mathrm{d}\sigma_{NLO}}{\mathrm{d}(\cos\theta_e)} = \left(\frac{G_F V_{ud}}{\sqrt{2}}\right)^2 f_{NLO}(E,\beta,\theta_e) \left[(1+3g_A^2)\mathcal{G}_1(\beta) + (1-g_A^2)\beta \mathcal{G}_2(\beta)\cos\theta_e \right]$$

Phase-space factor:

$$f_{NLO}(E, \beta, \theta_e) = f_{LO}(E, \beta) \left[1 - \frac{E}{m_N} \left(1 - \frac{E_{\nu}}{\beta E} \cos \theta_e \right) \right] + \mathcal{O}(m_N^{-2})$$

Correction factors:

$$\mathcal{G}_i(\beta) = 1 + \frac{\alpha}{2\pi} \mathcal{G}_i^{rad}(\beta) + \frac{1}{m_N} \mathcal{G}_i^{recoil}(\beta) \quad i = 1, 2$$

The factors $G_i(\beta)^{recoil}$ contain the "weak" magnetism interactions.

The pion-pole in the axial current and radiative corrections to the recoil terms are ignored

$$\left(\frac{Q}{m_{\pi}}\right)^2 \simeq 3 \cdot 10^{-5}$$
 and $\left(\frac{\alpha}{2\pi}\right) \mu_V \frac{Q}{2m_N} \simeq 10^{-6}$

We use the standard definition of G_F , given by the purely leptonic muon decay

$$\mu^- \to e^- + \bar{\nu}_e + \nu_{\mu}$$
.

The electro-weak $W^ Z^0$ box diagram is included in G_F , i.e. G_F includes the factor:

$$1 + \frac{3\alpha}{4\pi} \ln \left(\frac{m_W}{m_Z} \right)$$

The LEC, e_V^R , represents the "inner" (short-distance) corrections. A rough estimate, which is consistent with Marciano and Sirlin (1986),

$$\frac{\alpha}{2\pi}e_V^R(\mu=m_N)\simeq 2\cdot 10^{-2}$$

The dominant part of these "inner" corrections are due to the eletro-weak loops.

In the standard theory approach Sirlin and Marciano (2006), Fukugita and Kubota (2004) and Vogel(1984) have derived these corrections. The EFT approach allows a systematic improvement where, once e_V^R is determined, all low-energy corrections are uniquely determined.

We now compare neutron β -decay and $\bar{\nu}_e + p \rightarrow e^+ + n$.

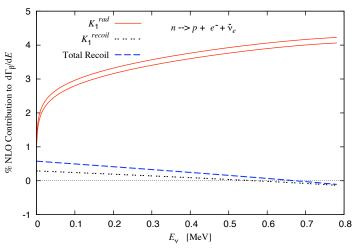
Neutron β -decay recoil corrections (Ando et al. (2004):

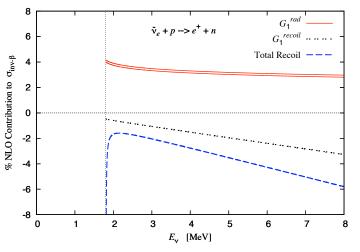
$$\mathcal{K}_{1}^{recoil} = \left(\frac{\beta^{2} E}{1 + 3g_{A}^{2}}\right) \left[1 + 2g_{A}\mu_{V} + g_{A}^{2}\right] + \left(\frac{E_{\nu}}{1 + 3g_{A}^{2}}\right) \left[1 - 2g_{A}\mu_{V} + g_{A}^{2}\right] + \mathcal{O}(m_{N}^{-1})$$
where we used $E^{max} = E_{\nu} + E + \mathcal{O}(m_{N}^{-1})$.

Recoil corrections for $\bar{\nu}_e + p \rightarrow e^+ + n$

$$\mathcal{G}_{1}^{recoil} = \left(\frac{\beta^{2} E}{1 + 3g_{A}^{2}}\right) \left[1 - 2g_{A}\mu_{V} + g_{A}^{2}\right] - \left(\frac{E_{\nu}}{1 + 3g_{A}^{2}}\right) \left[1 + 2g_{A}\mu_{V} + g_{A}^{2}\right]$$

Observe the sign changes for the $2g_A\mu_V$ terms, which are the dominant ones.





The NLO total cross section

$$\sigma = (G_F V_{ud})^2 \frac{\tilde{E}^2 \tilde{\beta}}{\pi} (1 + 3g_A^2) \left(1 + \frac{\alpha}{2\pi} \mathcal{G}_1^{rad}(\beta) \right)
\times \left\{ 1 + \frac{1}{m_N} \left[\mathcal{G}_1^{recoil}(\beta) - \tilde{E} - \left(\frac{1 + \tilde{\beta}^2}{\tilde{\beta}^2} \right) \left(E_{\nu} + \frac{(m_n - m_p)^2 - m_e^2}{2\tilde{E}} \right) \right.
\left. + \left(\frac{1 - g_A^2}{1 + 3g_A^2} \right) E_{\nu} \right] \right\}$$

These EFT recoil corrections agree with what was derived by Vogel and Beacom (1999).

Here
$$\tilde{E}=E_{\nu}-(m_n-m_p)=E+\mathcal{O}(m_N^{-1})$$
 and $\tilde{\beta}=\sqrt{\tilde{E}^2-m_e^2}/\tilde{E}=\beta+\mathcal{O}(m_N^{-1})$,

i.e. \tilde{E} and $\tilde{\beta}$ ignore the neutron recoil.

As seen earlier the NLO phase-space factor ($E=E_{\nu}+m_{p}-E_{n}$)

$$f_{NLO}(E, \beta, \theta_e) = \frac{E^2 \beta}{\pi} \left[1 - \frac{E}{m_N} \left(1 - \frac{E_\nu}{\beta E} \cos \theta_e \right) \right] + \mathcal{O}(m_N^{-2})$$

has no singularity as the positron velocity $\beta \to 0$.

CONCLUSIONS

- The estimated value for the radiative LEC $e_V^R(m_N)$ is dominated by electro-weak physics.
- The recoil corrections (m_N^{-1}) are very small (< 1%) in neutron β -decay.
- The recoil corrections for $\bar{\nu}_e + p \rightarrow e^+ + n$ are significant (a few %) and as large as the radiative corrections.

 The energy dependences are very different for these two corrections.
- ullet To extract an accurate value for the neutrino oscillating angle Θ_{13} from $ar{
 u}_e+p o e^++n$ require attention to the energy dependences of the two corrections which are of opposite signs.